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action
We suggest the following gauge fixed action in the classical limit [1]:

$$\Gamma^{(0)} = \Gamma_{\text{inv}} + \Gamma_m + \Gamma_{\text{gf}},$$

$$\Gamma_{\text{inv}} = \frac{1}{4} \int d^4x F_{\mu\nu} \star F_{\mu\nu},$$

$$\Gamma_m = \frac{\Omega^2}{4} \int d^4x \left(\frac{1}{2} \{ \tilde{x}_\mu ; A_\nu \} \star \{ \tilde{x}_\mu ; A_\nu \} + \{ \tilde{x}_\mu ; \bar{c} \} \star \{ \tilde{x}_\mu ; c \} \right) =$$

$$= \frac{\Omega^2}{8} \int d^4x (\tilde{x}_\mu \star c_\mu),$$

$$\Gamma_{\text{gf}} = \int d^4x \left[B \star \partial_\mu A_\mu - \frac{1}{2} B \star B - \bar{c} \star \partial_\mu s A_\mu - \frac{\Omega^2}{8} \bar{c}_\mu \star s c_\mu \right]$$

with

$$c_\mu = \left(\{ \tilde{x}_\mu ; A_\nu \} ; A_\nu \right) + \{ \tilde{x}_\mu ; \bar{c} \} ; c + \{ \bar{c} ; \tilde{x}_\mu ; c \}$$

and

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu - ig [A_\mu ; A_\nu],$$

$$\tilde{x}_\mu = (\theta^{-1})_{\mu\nu} x_\nu,$$

$$i\theta_{\mu\nu} = [x_\mu ; x_\nu].$$

BRST symmetry

The action is invariant under the BRST transformations given by

$$sA_\mu = D_\mu c = \partial_\mu c - ig [A_\mu ; c], \quad s\bar{c} = B,$$

$$sc = igc \star c, \quad sB = 0,$$

$$s\tilde{x}_\mu = \tilde{x}_\mu, \quad s^2 \varphi = 0 \quad \forall \varphi \in \{A_\mu, B, c, \bar{c}, \tilde{c}_\mu\}.$$

Furthermore, the fermionic multiplier field \tilde{c}_μ imposes an additional constraint, namely on-shell BRST invariance of c_μ and hence of the mass terms Γ_m .

Feynman rules

The bilinear parts of the action lead to the following improved propagators:

$$G_{\mu\nu}^A(x-y) = (-\Delta_4 + \Omega^2 \tilde{x}^2)^{-1} \delta_{\mu\nu} \delta^4(x-y),$$

$$G^{c\bar{c}}(x-y) = (-\Delta_4 + \Omega^2 \tilde{x}^2)^{-1} \delta^4(x-y),$$

$$G_\mu^{BA}(x-y) = (-\Delta_4 + \Omega^2 \tilde{x}^2)^{-1} \partial_\mu \delta^4(x-y),$$

$$G^B(x-y) = \left[\partial_\mu (-\Delta_4 + \Omega^2 \tilde{x}^2)^{-1} \partial_\mu - 1 \right] \delta^4(x-y).$$

Both the gauge field and the ghost propagators are essentially the Mehler kernel, which in momentum space reads

$$K_M(p, q) = \frac{\omega^3}{8\pi^2} \int_0^\infty d\alpha \frac{1}{\sinh^2(\alpha)} \exp\left(-\frac{\omega}{4} u^2 \coth \frac{\alpha}{2} - \frac{\omega}{4} v^2 \tanh \frac{\alpha}{2}\right),$$

$$u = (p-q), \quad v = (p+q),$$

so we may expect improved IR behaviour of the Feynman graphs. Since there are no vertices involving the B field and since the additional multiplier \tilde{c}_μ has no propagator, neither field will play a role in loop corrections.

The important vertices of the model in momentum space are given by

$$\tilde{V}_{\rho\sigma\tau}^{3A}(k_1, k_2, k_3) = 2ig(2\pi)^4 \delta^4(k_1 + k_2 + k_3) [(k_3 - k_2)_{\rho\sigma\tau} + (k_1 - k_3)_{\sigma\rho\tau} + (k_2 - k_1)_{\tau\rho\sigma}] \sin\left(\frac{k_1 \times k_2}{2}\right),$$

$$\tilde{V}_{\rho\sigma\tau}^{4A}(k_1, k_2, k_3, k_4) = -4g^2(2\pi)^4 \delta^4(k_1 + k_2 + k_3 + k_4) \times \left[(\delta_{\rho\sigma}\delta_{\tau\epsilon} - \delta_{\rho\epsilon}\delta_{\sigma\tau}) \sin\left(\frac{k_1 \times k_2}{2}\right) \sin\left(\frac{k_3 \times k_4}{2}\right) + (\delta_{\rho\sigma}\delta_{\tau\epsilon} - \delta_{\rho\epsilon}\delta_{\sigma\tau}) \sin\left(\frac{k_1 \times k_3}{2}\right) \sin\left(\frac{k_2 \times k_4}{2}\right) + (\delta_{\rho\sigma}\delta_{\tau\epsilon} - \delta_{\rho\epsilon}\delta_{\sigma\tau}) \sin\left(\frac{k_2 \times k_3}{2}\right) \sin\left(\frac{k_1 \times k_4}{2}\right) \right],$$

$$\tilde{V}_{\rho\sigma}^c(q_1, q_2, q_3) = -2ig(2\pi)^4 \delta^4(q_1 + q_2 + q_3) q_{3\rho} \sin\left(\frac{q_1 \times q_3}{2}\right),$$

equations of motion

$$\frac{\delta\Gamma^{(0)}}{\delta B} = \partial_\mu A_\mu - B + \frac{\Omega^2}{8} \left(\{ \tilde{x}_\mu ; c \} ; \tilde{c}_\mu - \{ \tilde{x}_\mu ; [\tilde{c}_\mu ; c] \} \right) = 0$$

$$\frac{\delta\Gamma^{(0)}}{\delta A_\nu} = (-\Delta_4 + \Omega^2 \tilde{x}^2) A_\nu + ig \left([A_\mu ; F_{\mu\nu}] + \partial_\mu [A_\mu ; A_\nu] + \{ \partial_\mu \bar{c} ; c \} \right) + \partial_\nu (\partial A) - \partial_\nu B + \frac{\Omega^2}{8} \left(\{ [D_\nu c ; \tilde{c}_\mu] ; \tilde{x}_\mu \} + \{ [D_\nu c ; \tilde{x}_\mu] ; \tilde{c}_\mu \} \right) - ig \frac{\Omega^2}{8} \left(\{ c ; \{ \tilde{x}_\mu ; \{ A_\nu ; \tilde{c}_\mu \} \} \} + \{ c ; \{ \tilde{c}_\mu ; \{ \tilde{x}_\mu ; A_\nu \} \} \} \right) = 0$$

$$\frac{\delta\Gamma^{(0)}}{\delta \bar{c}} = (-\Delta_4 + \Omega^2 \tilde{x}^2) \bar{c} + ig \partial_\mu [A_\mu ; c] - ig \frac{\Omega^2}{8} \left(\{ \tilde{x}_\mu ; c \star c \} ; \tilde{c}_\mu + \{ \tilde{x}_\mu ; \{ \tilde{c}_\mu ; c \star c \} \} \right) = 0$$

$$\frac{\delta\Gamma^{(0)}}{\delta c} = (\Delta_4 - \Omega^2 \tilde{x}^2) c + \frac{\Omega^2}{8} \left(\{ \tilde{c}_\mu ; \{ \tilde{x}_\mu ; B \} \} + \{ \tilde{x}_\mu ; \{ \tilde{c}_\mu ; B \} \} \right) - ig [A_\mu ; \partial_\mu \bar{c}] - \frac{\Omega^2}{8} D_\nu \left(\{ \tilde{x}_\mu ; \{ A_\nu ; \tilde{c}_\mu \} \} + \{ \{ \tilde{x}_\mu ; A_\nu \} ; \tilde{c}_\mu \} \right) + ig \frac{\Omega^2}{8} \left([c ; [\tilde{c}_\mu ; \{ \tilde{x}_\mu ; \bar{c} \}]] - [c ; \{ \tilde{x}_\mu ; [\bar{c} ; \tilde{c}_\mu] \}] \right) = 0$$

$$\frac{\delta\Gamma^{(0)}}{\delta \tilde{c}_\mu} = -\frac{\Omega^2}{8} s c_\mu = 0$$

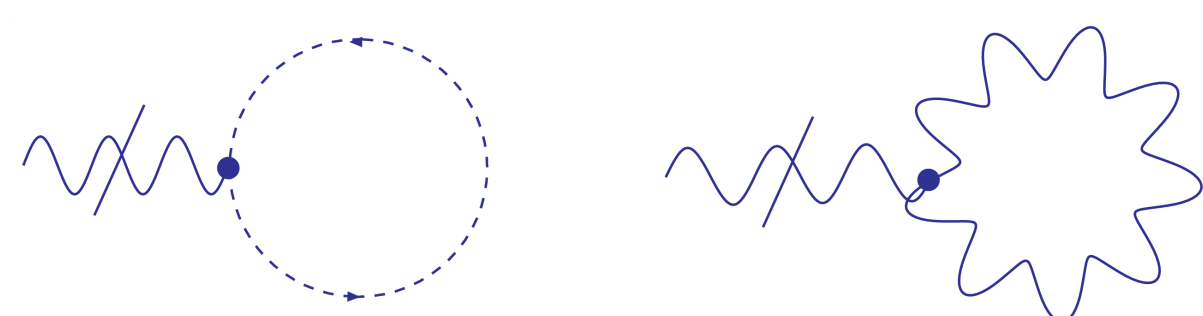
properties of the star product

$$A_1(x) \star A_2(x) = e^{\frac{i}{2} \theta^{\mu\nu} \partial_\mu \partial_\nu} A_1(x) A_2(x) \Big|_{x=y},$$

$$\int d^4x A_\mu(x) \star A_\nu(x) \star A_\rho(x) = \int d^4x A_\mu(x) \star A_\rho(x) \star A_\nu(x),$$

$$\int d^4x A_\mu(x) \star A_\nu(x) = \int d^4x A_\nu(x) \star A_\mu(x),$$

$$\{ \tilde{x}_\mu ; A_\nu(x) \} = 2\tilde{x}_\mu A_\nu(x).$$



gauge field tadpole with amputated external leg

• $\Omega \neq 0$

$$T_\rho = \int d^4k \int d^4k' \tilde{K}_M(k, k') V_{\rho\sigma\tau}^{3A}(p, -k', k) \delta_{\sigma\tau}$$

$$= 6i\pi^4 g \tilde{p}_\rho \int d\alpha \frac{\sinh \alpha}{(\cosh \alpha - 1)^3}$$

$$\times \exp\left(\frac{\omega}{4} \beta^2 \frac{\cosh \alpha - 1}{\sinh \alpha} - \beta^2 \frac{\theta^2}{16} \frac{\sinh \alpha}{\cosh \alpha - 1}\right)$$

$$= \text{finite}$$

• $\Omega = 0$

$$T_\rho = \sum_{\eta=\pm 1} i\eta^2 3(2\pi)^4 \pi^2 g \left(\delta^{(4)}(\rho) \tilde{p}_\rho \right) \int_0^\infty d\alpha \frac{1}{\alpha^3} \exp\left\{-\frac{\beta^2}{4\alpha}\right\}$$

$$= 6i(2\pi)^4 \pi^2 g \left(\delta^{(4)}(\rho) \tilde{p}_\rho \right) \frac{16}{\beta^4} \rightarrow \infty$$

open questions

- The bilinear parts of the action are Langmann-Szabo invariant, but some vertices are not and the question remains whether this is sufficient to remove UV/IR mixing problems. Loop calculations are currently work in progress.
- What are the (classical) solutions to the equations of motion?
- Is there a connection between this model and induced gauge theory [2,3].

references

- [1] D. N. Blaschke, H. Grosse and M. Schweda, *Non-Commutative $U(1)$ Gauge Theory on R^{**4} with Oscillator Term and BRST Symmetry*, *Eur. Phys. Lett.* **79** (2007) 61002, [arXiv:0705.4205].
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