

Towards renormalizable models for gauge fields in non-commutative space



Background:

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \langle T_{\mu\nu} \rangle$$
,

$$\Delta x_{\mu} \simeq \text{ Planck length } \lambda_{\rho} = \sqrt{\frac{G\hbar}{c^3}} \simeq 10^{-33} cm^{-3}$$

 $\Pi^{n \text{ npl-ins.}}(p) \equiv \lambda^2 \int d^4k \, \frac{\mathrm{e}^{\mathrm{i}k\tilde{p}}}{\left(\tilde{k}^2\right)^n \left[k^2 + m^2 + \frac{\mathrm{e}^2}{\tilde{k}^2}\right]^{n+1}}$



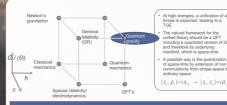
$[x^{\mu},x^{\nu}]=i\Theta^{\mu}$

- istory

 H. Inyder (1946/47): "minimal length" to smear out point-like interactions as UV regularization in QFT

 early 1990s: J. Madore, A. Connes, T. Filk and others
 (fuzzy sphere, estensive study of non-commutative geometry, ne divergences UV/IR mixing)

 N. Seberg & E. Witten (1999): connection to string theory (effective low-energy action on D-branes with strong B-field background, non-commutativity parameter related to B-field as \$\frac{a}{a} \cdots \frac{a}{a} \cdots \frac{a}{a



$$_{i}^{\mu},\hat{\chi}^{\nu}]=\mathrm{i}\theta^{\mu\nu}$$
 \Rightarrow implies uncertainty relation
$$\Delta\chi^{\mu}\Delta\chi^{\nu}\geq\frac{1}{2}|\theta^{\mu\nu}|\sim(\lambda_{p})^{2}\,,$$

ere $\lambda_{\rm o} \approx 10^-33{\rm cm}$ is the Planck length • definition of the Groenewold-Moyal *-product

$$f(x) \star g(x) = e^{\frac{1}{2}\theta^{\mu\nu}\partial_{\mu}^{x}\partial_{\nu}^{y}} f(x)g(y)\Big|_{x=y}$$

$$\neq g(x) \star f(x)$$

ordinates x instead of operators \hat{x} , since

$$[x^{\mu} , x^{\nu}] = i\theta^{\mu\nu}$$

$$\begin{split} \int \mathrm{d}^4 x \, f(x) \star g(x) \star h(x) &= \int \mathrm{d}^4 x \, h(x) \star f(x) \star g(x) \\ \Longrightarrow \int \mathrm{d}^4 x f(x) \star g(x) &= \int \mathrm{d}^4 x f(x) g(x) \end{split}$$

$$\int d^4k \frac{e^{ik\tilde{p}}}{k^2+i\epsilon} \propto \frac{1}{\tilde{p}^2} \quad \text{with} \quad \tilde{p}^\mu = \theta^{\mu\nu} p_\nu$$

⇒ origin of the UV/IR mixing p



$$S_{\phi 4} = \int d^4x \left[\frac{1}{2} \left(\partial^{\mu}\phi \star \partial_{\mu}\phi + m^2\phi \star \phi \right) + \frac{\lambda}{4!} \phi^{\star 4} \right],$$

 $S_{YM} = -\frac{1}{4} \int d^4x F_{\mu\nu} \star F^{\mu\nu} + \dots$

with $F_{\mu\nu}=\partial_{\mu}A_{\nu}-\partial_{\nu}A_{\mu}-ig\left[A_{\mu}\stackrel{*}{,}A_{\nu}\right]$ ead to (gauge independent) IR singular self-energy graphs

$$\begin{split} &\Pi^{\mu\nu}_{\text{OA,IR}}(\rho) \propto \frac{1}{\tilde{p}^2}\,, \\ &\Pi^{\mu\nu}_{\text{YM,IR}}(\rho) \propto \frac{\tilde{p}^\mu \tilde{p}^\nu}{(\tilde{p}^2)^2} \quad \text{with} \quad \tilde{p}^\mu = \theta^{\mu\nu} \rho_\nu \\ &\Rightarrow \text{Graphs with these insertions are IR div} \end{split}$$



$$S = \int d^4x \left[\frac{1}{2} \left(\partial^{\mu}\phi \star \partial_{\mu}\phi + m^2\phi \star \phi - \phi \star \frac{a^2}{\Box}\phi \right) + \frac{\lambda}{4!} \phi^{\star 4} \right]$$

ator: $G(k) = \frac{1}{k^2 + m^2 + \frac{s^2}{k^2}} = \frac{1}{2} \sum_{\zeta = \pm 1} \frac{1 + \zeta \frac{m^2}{2M^2}}{k^2 + \frac{m^2}{2} + \zeta M^2}$

 \Rightarrow new "damping" behaviour $\lim_{k \to 0} G(k) = 0$





Dimensions of the Universe:

Physically accessible

"Terra incognita"

Gauge fields:

$$\int \mathrm{d}^4 x \phi(x) \frac{a^2}{\theta^2 \square} \phi(x) \quad \Rightarrow \quad \int \mathrm{d}^4 x \frac{1}{4} F^{\mu\nu} \star \frac{a^2}{D^2 \widetilde{D}^2} \star F_{\mu\nu}$$

$$\begin{split} &U_{\star}(1):\delta_{\varepsilon}A_{\mu}=\partial_{\mu}\varepsilon+\mathrm{i}g\left[A_{\mu}\stackrel{\star}{,}\varepsilon\right]\,,\\ &F_{\mu\nu}=\partial_{\mu}A_{\nu}-\partial_{\nu}A_{\mu}-\mathrm{i}g\left[A_{\mu}\stackrel{\star}{,}A_{\nu}\right]\,,\\ &D_{\mu}\bullet\equiv\partial_{\mu}\bullet-\mathrm{i}g\left[A_{\mu}\stackrel{\star}{,}\bullet\right]\quad\text{and}\quad \widetilde{D}_{\mu}=\theta_{\mu\nu}D^{\nu} \end{split}$$
on $\frac{1}{D^2}F\equiv Y$ transforms covariantly $(sY=ig\ [c\ ;\ Y])$ and is to stood as formal power series in the gauge field A_μ



Since
$$=\int \!\mathrm{d}^4x \left[\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{\lambda}{2} \left(B_{\mu\nu} + \bar{B}_{\mu\nu} \right) F^{\mu\nu} - \right. \\ \left. - \mu^2 \bar{B}_{\mu\nu} D^2 \widetilde{D}^2 B^{\mu\nu} + \mu^2 \bar{\psi}_{\mu\nu} D^2 \widetilde{D}^2 \psi^{\mu\nu} \right]$$

$$S_{gf} = \int d^4x \left[b \partial_\mu A^\mu - \bar{c} \partial^\mu D_\mu c \right]$$

agator with IR damping:

$$G_{\mu\nu}^{AA}(k) = rac{1}{\left(k^2 + rac{\lambda^2}{\mu^2 k^2}
ight)} \left(g_{\mu
u} - rac{k_\mu k_
u}{k^2}
ight)$$

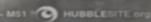
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 standard top-down renormalization schemes fail to succeed for non-commutative theories! Scalar non-commutative models were proved to be renormalizable using schemes which would break gauge invariance in our model.

 Therefore, we need to develop new techniques . . .

 $S_{break} = \int d^4x \frac{\lambda}{2} B_{\mu\nu} \star F^{\mu\nu}$ \Rightarrow $\int d^4x \, s \left(\bar{Q}_{\mu\nu\alpha\beta} \star B^{\mu\nu} \star F^{\alpha\beta} \right)$





 $sF_{\mu\nu} = ig [c \uparrow F_{\mu\nu}]$, $s\bar{\psi}_{\mu\nu} = \bar{B}_{\mu\nu} + ig \left\{c \uparrow \bar{\psi}_{\mu\nu}\right\}$ $sB_{\mu\nu} = \psi_{\mu\nu} + ig \left[c \uparrow B_{\mu\nu}\right],$